

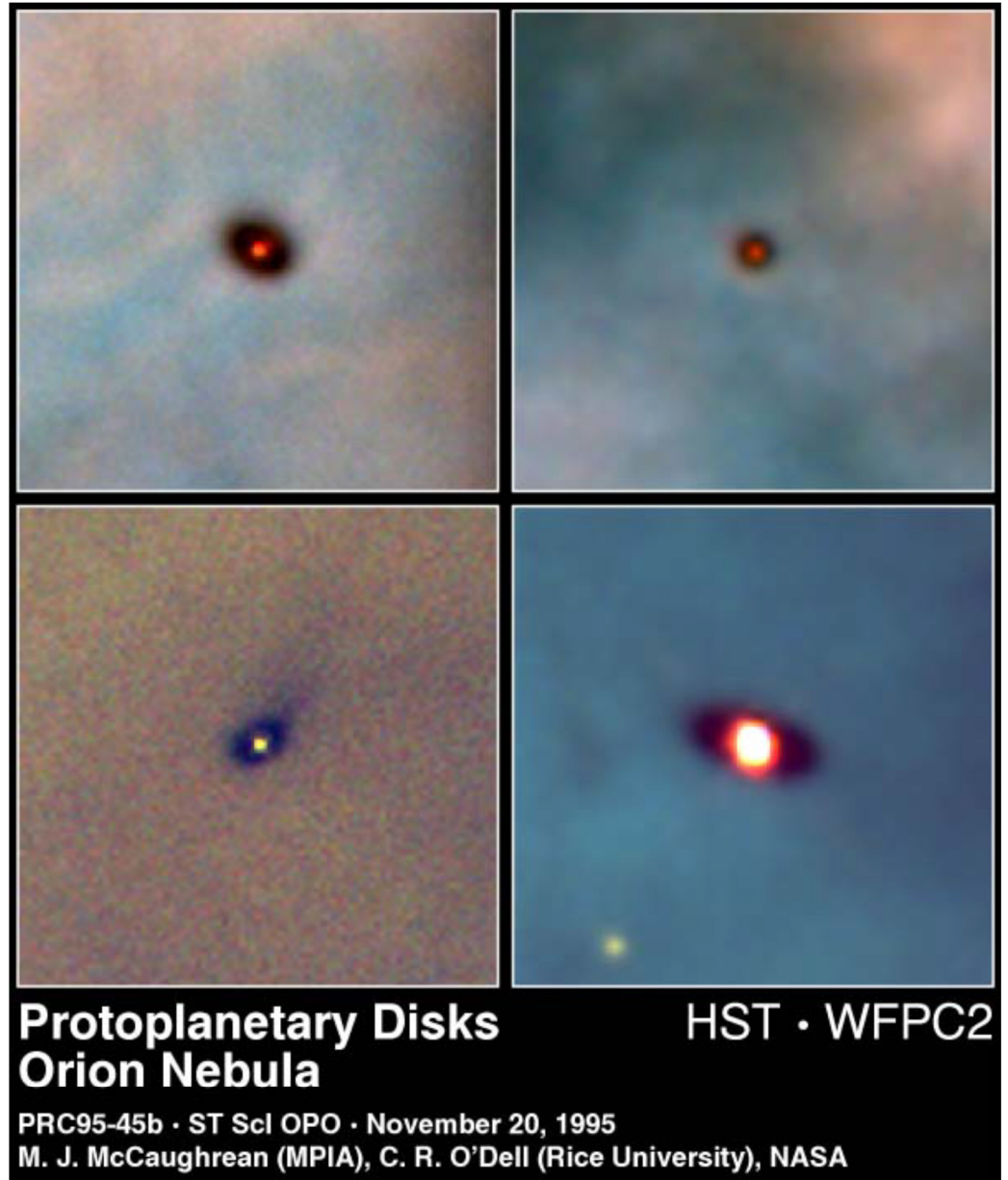
# How Did the Earth Form?



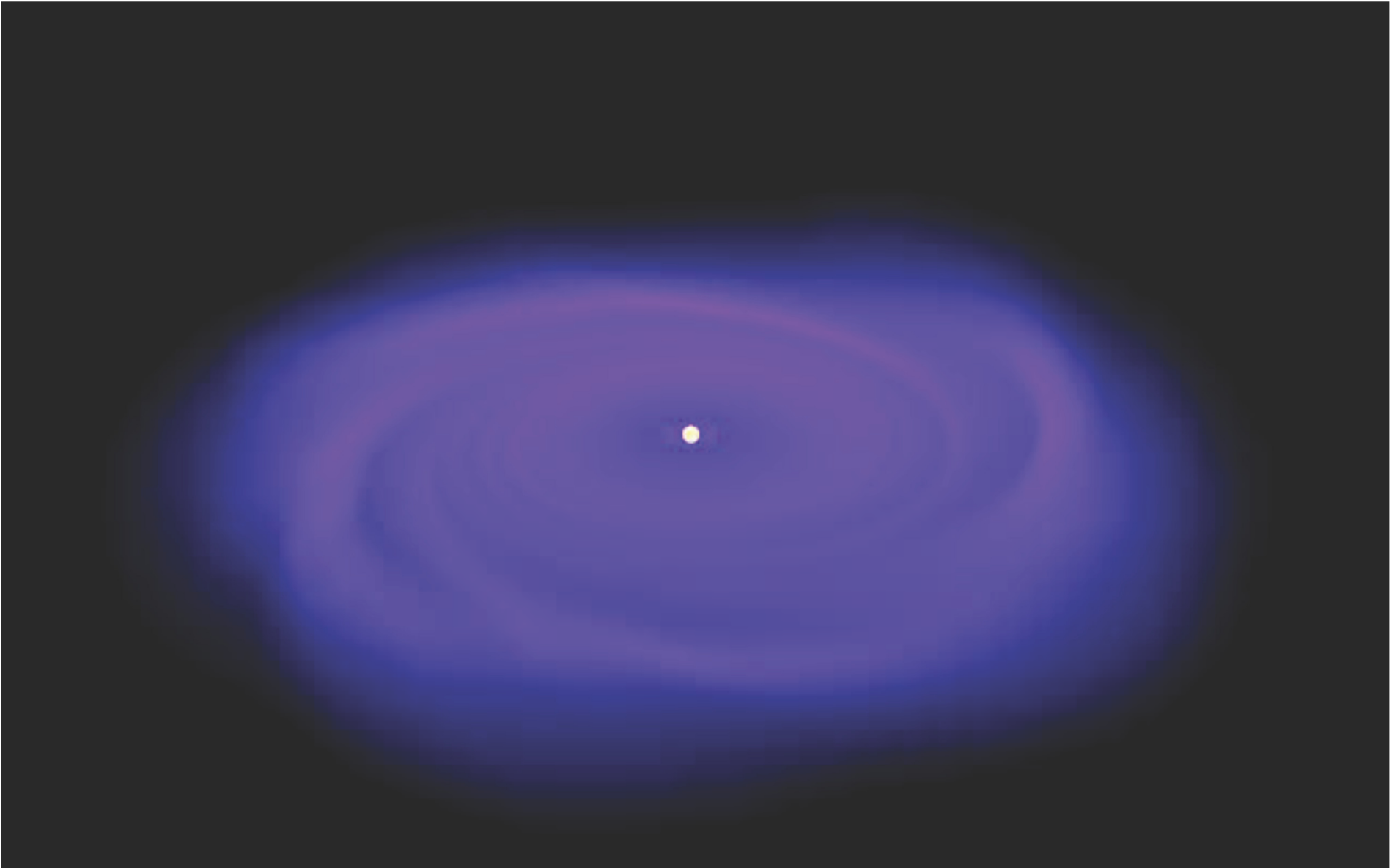
Joseph Lazio  
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# Proto-planetary Disks

- What we know (or think we know)
  - All of the planets in the solar system nearly in a plane.
  - Disks of dust and gas detected around nearby forming stars.
  - Simulations of proto-planetary disks show planets growing.
- Ergo, planets form in disks of material around stars.



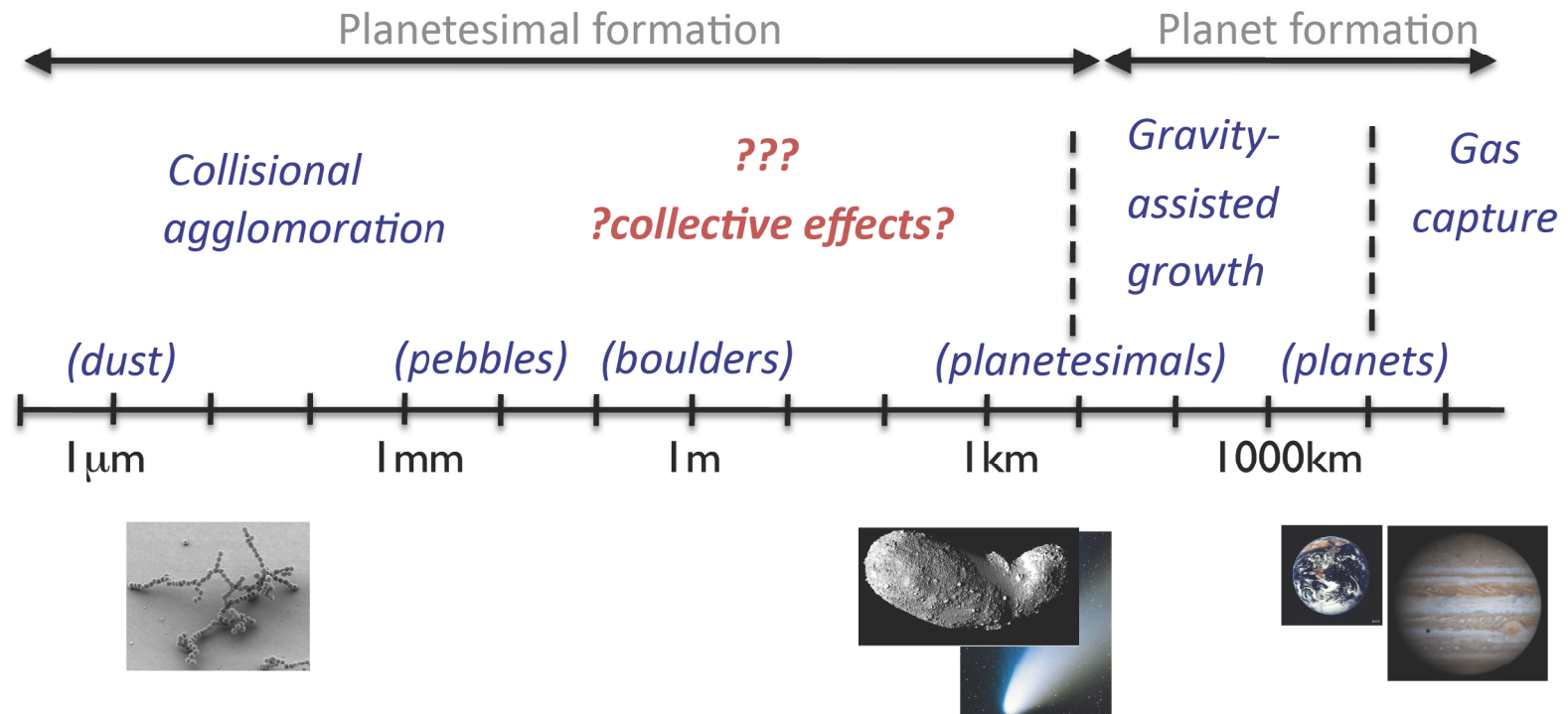
# Simulation of Planet Formation



# Proto-planetary Disks

## From Dust to Planets

(what we think happens ...)

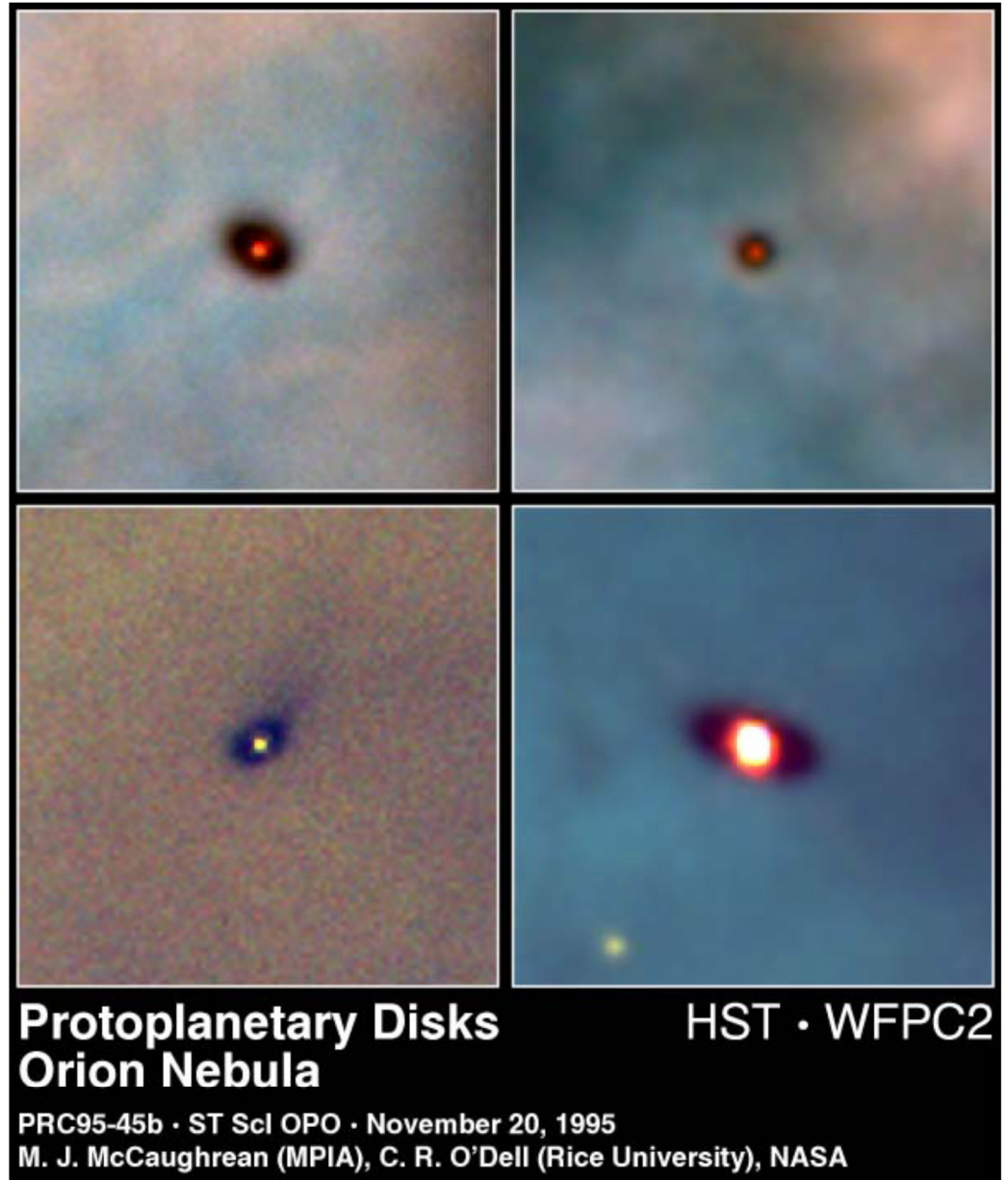


(From D. Wilner)



# Proto-planetary Disks

- *All of the planets in the solar system nearly in a plane.*
  - *Disks of dust and gas detected around nearby forming stars.*
  - *Ergo, planets form in disks of material around stars.*
  - Want to understand how planets grow in proto-planetary disks
- Disks are **opaque** to visible light.
- Want **observations at 5 GHz to 1 THz** → mm- and cm-wavelength radiation, comparable to size of particles.



Protoplanetary Disks  
Orion Nebula

HST · WFPC2

PRC95-45b · ST ScI OPO · November 20, 1995

M. J. McCaughrean (MPIA), C. R. O'Dell (Rice University), NASA

# Radio Astronomy Service and Planet Formation

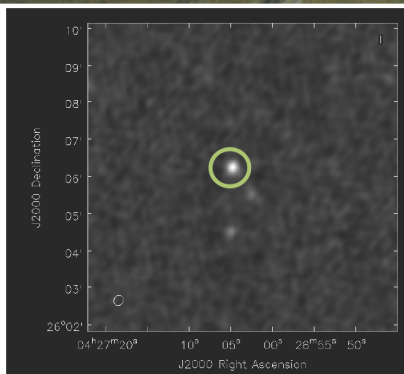
## Expanded Very Large Array

(1 GHz to 50 GHz; US-Canada-Mexican partnership)



## Atacama Large Millimeter/submillimeter Array

(100 GHz to 1 THz; US-European-Japanese partnership)



**EVLA observations:** ongoing—60+ nearby stars with disks  
(example shows C-band observations of disk around star DG Taurus, illustrating decimeter-sized particles in the disk)

**ALMA status:** in commissioning





# Formation of Planets and the Radio Spectrum



Understanding the formation of planets, and how we came to be, requires

- Infrared telescopes in space;  
... and X-ray telescopes in space?
  - **E-S**: 2110–2120 MHz, 7145–7190 MHz, 34.2–34.7 GHz
  - **S-E**: 2290–2300 MHz, 8400–8450 MHz, 31.8–32.3 GHz
- Telescopes on the ground operating from 5 GHz to 1 THz.

